

Smooth Imploding Solutions for 3D Compressible Fluids

T. Buckmaster (Princeton University & IAS)

with Cao-Labora (MIT) and Gómez-Serrano (Brown & Univ. Barcelona)

Long Time Behavior and Singularity Formation in PDEs-Part IV,
SITE online conference
January 18, 2022

Compressible Euler Equations

Non-Isentropic Form

Full **non-isentropic Euler equations**:

$$\partial_t(\rho \mathbf{u}) + \operatorname{div}(\rho \mathbf{u} \otimes \mathbf{u} + p \operatorname{Id}) = 0 \quad (\text{Conservation of momentum})$$

$$\partial_t \rho + \operatorname{div}(\rho \mathbf{u}) = 0 \quad (\text{Conservation of mass})$$

$$\partial_t E + \operatorname{div}((p + E)\mathbf{u}) = 0 \quad (\text{Conservation of energy})$$

where \mathbf{u} is the **velocity**, ρ , the **density**, p , the **pressure** and E , the **energy**. Conservation of energy can be replaced by **transport of specific entropy** $\partial_t \mathbf{S} + \mathbf{u} \cdot \nabla \mathbf{S} = 0$. The pressure is

$$p = (\gamma - 1)(E - \tfrac{1}{2}\rho|\mathbf{u}|^2) = \tfrac{1}{\gamma}\rho^\gamma e^{\mathbf{S}}$$

for adiabatic exponent $\gamma > 1$. The **sound speed** is given by

$$c = \sqrt{\frac{\gamma p}{\rho}}.$$

Shock waves and imploding solutions

- ▶ **Shock waves:** The prototypical singularity for the Euler equations is a shock wave, which occurs when the speed of a disturbance exceeds the local speed of sound. Mathematically, one is interested in both the formation of the shock and the development of the shock.
- ▶ **Implosions:** Implosions involve spherically symmetric solutions that collapse at a point in finite time. Classically, one considers imploding shock waves. Recently, Merle-Raphael-Rodnianski-Szeftel showed there exist smooth imploding solutions.

Shock formation results

- ▶ Christodoulou '07, Christodoulou-Miao '14: 3D isentropic, **irrotational**.
- ▶ Luk-Speck '18: 2D isentropic, **non-trivial vorticity**.
- ▶ B-Shkoller-Vicol '19: 2D isentropic, azimuthal, **non-trivial vorticity** + **description of self-similar profile**.
- ▶ B-Shkoller-Vicol '19: 3D isentropic, **non-trivial vorticity** + **description of self-similar profile**.
- ▶ B-Shkoller-Vicol '20: **Full 3D Euler** + **description of self-similar profile**.
- ▶ Luk-Speck '21: **Full 3D Euler**, allow non-generic shocks.

Related works: John-Klainerman '84, Klainerman, John '87, Hörmander '87, John '81, Sideris '85, Alinhac '99

Shock development results

- ▶ [Lebaud '94](#): 1D, 2x2 p-system, existence of discontinuous shock (uniqueness follows by T.P. Liu) (cf. Chen-Dong '01, Kong '02 for generalizations)
- ▶ [Yin '04](#): Spherically symmetric Euler, existence of weak solution past formation, **no uniqueness and no description of weak discontinuities**.
- ▶ [Christodoulou-Lisbach '16](#): Spherically symmetric isentropic Euler in formation, restricted problem thereafter, **not a weak solution to Euler**.
- ▶ [Christodoulou '19](#): Multi-D, irrotational, isentropic shock development for the restricted problem, **not a weak solution to Euler**.
- ▶ [B-Drivas-Shkoller-Vicol '21](#): Development for full Euler under azimuthal symmetry satisfying the Rankine-Hugoniot jump conditions, **uniqueness, full description of weak discontinuities**.

Implosion results

- ▶ [Guderley '42](#): Self-similar imploding **shock waves** solutions to Euler.
- ▶ [Merle-Raphael-Rodnianski-Szeftel '19](#): Smooth imploding self-similar solutions exist from **a.e.** adiabatic exponent $\gamma > 1$.
- ▶ [Biasi' 21](#): Detailed numerical description of smooth self-similar imploding solutions.

Related work:

Navier-Stokes (Merle et al. '19), NLS via Madelung transform (Merle et al. '19), Euler Poisson (Guo-Hadzic-Jang-Schrecker '21)

Setup

- Isentropic, spherically symmetric Euler

$$\partial_t u + u \partial_R u + \frac{1}{\gamma \rho} \partial_R \rho^\gamma = 0 \quad \text{and} \quad \partial_t \rho + \frac{1}{R^2} \partial_R (R^2 \rho u) = 0,$$

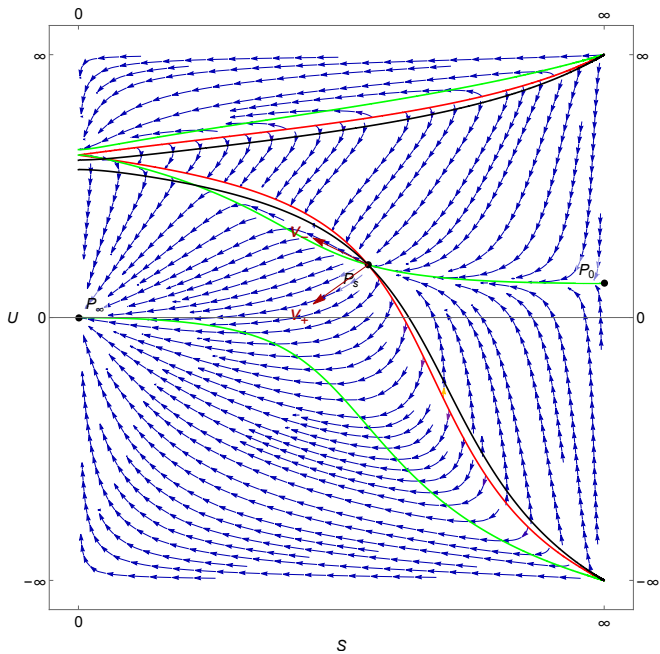
- The self-similar ansatz

$$u(r, t) = r^{-1} \frac{R}{T-t} U\left(\log\left(\frac{R}{(T-t)^{\frac{1}{\gamma}}}\right)\right) \quad \text{and} \quad \sigma(r, t) = \alpha^{-\frac{1}{2}} r^{-1} \frac{R}{T-t} S\left(\log\left(\frac{R}{(T-t)^{\frac{1}{\gamma}}}\right)\right),$$

where $\sigma = \frac{1}{\alpha} \rho^\alpha$ is the **rescaled sound speed**.

- Setting $\xi = \log\left(\frac{R}{(T-t)^{\frac{1}{\gamma}}}\right)$ leads to the autonomous ODE

$$\frac{dU}{d\xi} = \frac{N_U(U, S)}{D(U, S)}, \quad \text{and} \quad \frac{dS}{d\xi} = \frac{N_S(U, S)}{D(U, S)}.$$



Result of Merle et al. '19

For a.e. $\gamma > 1$, there exists a countably infinite sequence of self-similar solutions to isentropic Euler. The velocity and density blow up at the origin.

The condition on γ relates to the non-vanishing of an analytic function. The condition is not proven for any specific γ , but may be checked numerically. The case $\gamma = 5/3$ (monatomic gases) is specifically ruled out.

Compressible Navier-Stokes

Isentropic 3D compressible Navier-Stokes with constant viscosity:

$$\partial_t(\rho u) + \operatorname{div}(\rho u \otimes u) + \nabla p(\rho) - \mu_1 \Delta u - (\mu_1 + \mu_2) \nabla \operatorname{div} u = 0,$$

$$\partial_t \rho + \operatorname{div}(\rho u) = 0,$$

for $\mu_1 \geq 0$ and $2\mu_1 + \mu_2 \geq 0$.

Merle et al. '19: there exists imploding solutions to NS for a.e.

$1 < \gamma < \frac{2+\sqrt{3}}{\sqrt{3}}$ with decaying density.

Previously, Xin '98: blow up for initial data with compact density and

Rozanova '08: blow up for rapidly decaying density.

Problems left open

1. Do imploding solutions for Euler exist for all $\gamma > 1$?
2. Can one construct imploding solutions to the Navier-Stokes equation with initial density constant at infinity?

Main result

1. There exists smooth self-similar imploding solutions **for all $\gamma > 1$** .
2. For the case $\gamma = \frac{7}{5}$ (diatomic gas, e.g. oxygen, hydrogen, nitrogen) there exists a countably infinite sequence of imploding solutions.
3. Simplified proofs of linear stability and non-linear stability.
4. Asymptotically self-similar imploding solutions to NS for $\gamma = \frac{7}{5}$.
5. **First example** of initial data with density **constant at infinity** leading to blow up for NS.

Riemann invariants

- ▶ Riemann invariants: $w = u + \sigma$ and $z = u - \sigma$.
- ▶ Self-similar ansatz

$$w(R, t) = \frac{1}{r} \cdot \frac{R}{T-t} W(\xi) \quad \text{and} \quad z(R, t) = \frac{1}{r} \cdot \frac{R}{T-t} Z(\xi)$$

- ▶ Setting $\xi = \log\left(\frac{R}{(T-t)^{\frac{1}{r}}}\right)$ yields

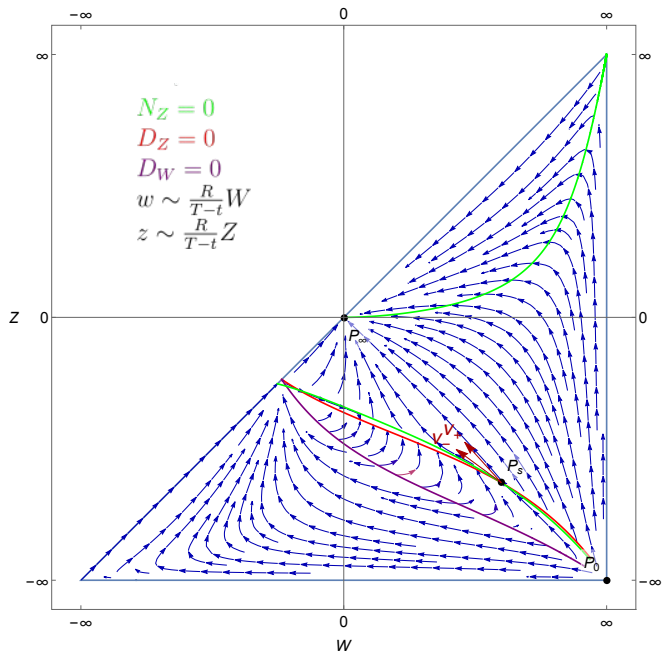
$$\left(r + \frac{1}{2}((1+2\alpha)W + (1-\alpha)Z)\right)W + \left(1 + \frac{1}{2}(W + Z + \alpha(W-Z))\right)\partial_{\xi}W - \frac{\alpha}{2}Z^2 = 0$$

$$\left(r + \frac{1}{2}((1-\alpha)W + (1+2\alpha)Z)\right)Z + \left(1 + \frac{1}{2}(W + Z - \alpha(W-Z))\right)\partial_{\xi}Z - \frac{\alpha}{2}W^2 = 0$$

- ▶ Rearranging,

$$\frac{dW}{d\xi} = \frac{-(r + \frac{1}{2}((1+2\alpha)W + (1-\alpha)Z))W + \frac{\alpha}{2}Z^2}{1 + \frac{1}{2}(W + Z + \alpha(W-Z))} = \frac{N_W}{D_W},$$

$$\frac{dZ}{d\xi} = \frac{-(r + \frac{1}{2}((1-\alpha)W + (1+2\alpha)Z))Z + \frac{\alpha}{2}W^2}{1 + \frac{1}{2}(W + Z - \alpha(W-Z))} = \frac{N_Z}{D_Z}.$$



Analysis of the point P_s

Under the change of variables $\xi \mapsto \psi$ where $\partial_\psi = -D_W D_Z \partial_\xi$:

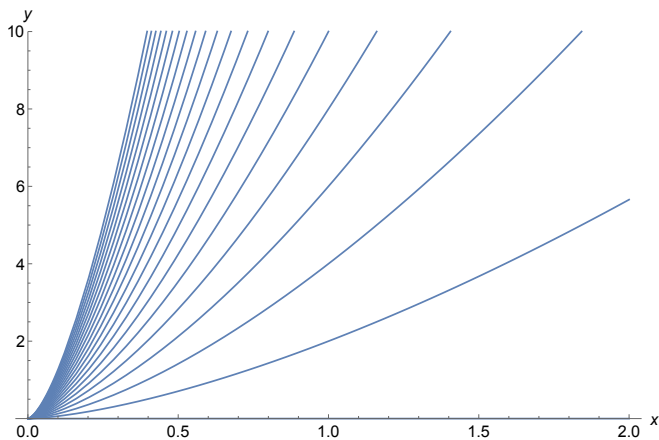
$$\partial_\psi W = -N_W D_Z \quad \text{and} \quad \partial_\psi Z = -N_Z D_W,$$

P_s becomes a **stable** stationary point. Consider the simple ODE:

$$\dot{x} = \lambda_+ x, \quad \dot{y} = \lambda_- y$$

for $\lambda_- < \lambda_+ < 0$. If $k = \frac{\lambda_-}{\lambda_+} \notin \mathbb{N}$, $x = 0$ and $y = 0$ are the sole smooth solutions. Non-smooth, C^k solutions exist of the form $y = Cx^k$ whose series agrees with the solution $y = 0$ up to order $\lfloor k \rfloor$.

Case $\lambda_- = -\frac{3}{2}$ and $\lambda_+ = -1$



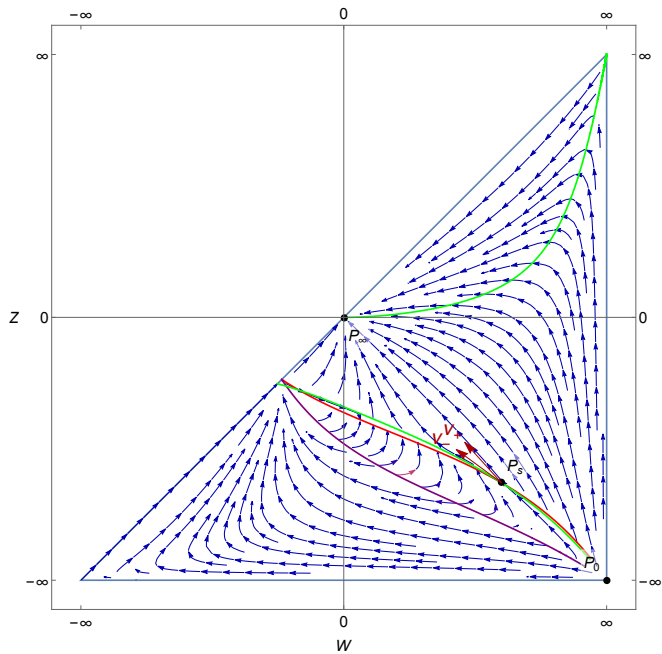
Returning to our ODE

$$\partial_\psi W = -N_W D_Z \quad \text{and} \quad \partial_\psi Z = -N_Z D_W,$$

Let $\lambda_- < \lambda_+ < 0$ be the eigenvalues of the Jacobian at P_s , and define

$$k = \frac{\lambda_-}{\lambda_+}.$$

If ν_- , ν_+ are the corresponding eigenvectors, we consider smooth solutions tangent to ν_- . The smooth solutions tangent to ν_+ correspond to the Guderley solutions.



Taylor Expansion around P_s ($\xi = 0$)

- Write the solution crossing P_s as a series

$$(W(\xi), Z(\xi)) = \sum_{n=0}^{\infty} \frac{\xi^n}{n!} (W_n, Z_n)$$

- If $D_{\circ,n} = \nabla D_{\circ} \cdot (W_n, Z_n)$ for $\circ \in \{W, Z\}$, $n \geq 1$, then

$$D_{W,0} W_n = N_{W,n-1} - \sum_{j=0}^{n-2} \binom{n-1}{j} D_{W,n-1-j} W_{j+1},$$

$$\begin{aligned} Z_n D_{Z,1} (n-k) &= - \sum_{j=1}^{n-2} \binom{n}{j} D_{Z,n-j} Z_{j+1} \\ &\quad + (N_{Z,n} - (\partial_Z N_Z(P_2)) Z_n) + Z_1 (-D_{Z,n} + Z_n \partial_Z D_Z(P_2)). \end{aligned}$$

The wiggle with r

- Restrict $1 < r < r^*$ where

$$r^*(\gamma) = \begin{cases} \frac{2}{\left(\sqrt{2}\sqrt{\frac{1}{\gamma-1}}+1\right)^2} + 1 & 1 < \gamma < \frac{5}{3}. \\ \frac{3\gamma-1}{2+\sqrt{3}(\gamma-1)} & \gamma \geq \frac{5}{3}. \end{cases}$$

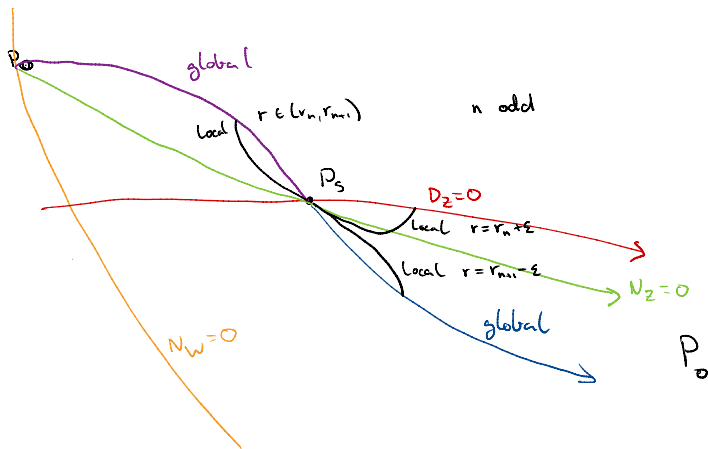
$k(r) : (1, r^*) \rightarrow \mathbb{R}$ is **increasing** with r and $k \rightarrow \infty$ as $r \rightarrow r^*$.

- For $j \in \mathbb{N}$, define r_j such that $j = k(r_j)$. At k_j the denominator of Z_n is singular and switches sign.
- To connect P_0 to P_∞ , we show that for n odd
 - For $r \in (r_n, r_{n+1})$ the solution to the **left** of P_c converges to P_∞ as $\xi \rightarrow \infty$.
 - For $r = r_n + \varepsilon$ the solution to the **right** of P_c intersects the line $D_W = 0$.
 - For $r = r_{n+1} - \varepsilon$ the solution to the **right** of P_c intersects the line $D_Z = 0$.

For $\gamma > 1$, this is shown for $n = 3$, for $\gamma = \frac{7}{5}$, this is shown for odd n sufficiently large.

Barriers

P_∞



The $\gamma = \frac{7}{5}$ case

- ▶ For $\gamma < \frac{5}{3}$ and fixed n , (W_n, Z_n) converge as $r \rightarrow r^*$ in a non-trivial manner.
- ▶ The *wiggle* can be determined from the sign and a lower bound on the coefficients of order $O(k)$.
- ▶ Via a computer assisted proof, compute the first 10000 coefficients with effective error bounds at r^* ($Z_{10000} \sim 6 \times 10^{46770}$).
- ▶ The proof works for general $\gamma < \frac{5}{3}$; however, the computation degenerates as $\gamma \rightarrow \frac{5}{3}$.

Strategy for stability for Euler and Navier-Stokes

The basic strategy is:

1. Linearize Euler/Navier-Stokes around self-similar profiles.
2. Show the linearized operator as finitely many unstable modes.
3. A Brouwer fix point argument shows there exists a manifold of initial data of finite co-dimension leading to imploding asymptotically self-similar blow-up.

Write

$$w(R, t) = r^{-1} (T - t)^{r^{-1}-1} \mathcal{W}\left(\frac{R}{(T-t)^{\frac{1}{r}}}, -\frac{\log(T-t)}{r}\right)$$

$$z(R, t) = r^{-1} (T - t)^{r^{-1}-1} \mathcal{Z}\left(\frac{R}{(T-t)^{\frac{1}{r}}}, -\frac{\log(T-t)}{r}\right).$$

and define the self-similar variables

$$s = -\frac{\log(T-t)}{r}, \quad \zeta = \frac{R}{(T-t)^{\frac{1}{r}}} = e^s R = \exp(\xi).$$

The Navier-Stokes equations ($\mu_1 = 1$ and $\mu_2 = -1$) become

$$\begin{aligned} & (\partial_s + r - 1) \mathcal{W} + \left(\zeta + \frac{1}{2}(\mathcal{W} + \mathcal{Z} + \alpha(\mathcal{W} - \mathcal{Z}))\right) \partial_\zeta \mathcal{W} + \frac{\alpha}{2\zeta} (\mathcal{W}^2 - \mathcal{Z}^2) \\ &= \frac{r^{1+\frac{1}{\alpha}} 2^{1/\alpha}}{\alpha^{1/\alpha} \zeta^2 ((\mathcal{W} - \mathcal{Z}))^{\frac{1}{\alpha}}} e^{(2-r+\frac{1}{\alpha}(1-r))s} \left(\partial_\zeta (\zeta^2 \partial_\zeta (\mathcal{W} + \mathcal{Z})) - 2(\mathcal{W} + \mathcal{Z})\right) \\ & (\partial_s + r - 1) \mathcal{Z} + \left(\zeta + \frac{1}{2}(\mathcal{W} + \mathcal{Z} - \alpha(\mathcal{W} - \mathcal{Z}))\right) \partial_\zeta \mathcal{Z} - \frac{\alpha}{2\zeta} (\mathcal{W}^2 - \mathcal{Z}^2) \\ &= \frac{r^{1+\frac{1}{\alpha}} 2^{1/\alpha}}{\alpha^{1/\alpha} \zeta^2 ((\mathcal{W} - \mathcal{Z}))^{\frac{1}{\alpha}}} e^{(2-r+\frac{1}{\alpha}(1-r))s} \left(\partial_\zeta (\zeta^2 \partial_\zeta (\mathcal{W} + \mathcal{Z})) - 2(\mathcal{W} + \mathcal{Z})\right) \end{aligned}$$

If $2 - r + \frac{1}{\alpha}(1 - r) < 0$ then the dissipation term is an exponentially decaying error term, i.e.

$$r > \frac{2\gamma}{\gamma + 1}, .$$

Recall $1 < r < r^*$. Then, the adiabatic exponents are restricted to

$$1 < \gamma < \frac{2+\sqrt{3}}{\sqrt{3}}$$

.

Linearization

Let (\bar{W}, \bar{Z}) be an exact, self-similar Euler profile and define $(\widetilde{W}, \widetilde{Z}) = (W - \bar{W}, Z - \bar{Z})$. Rewrite Euler/NS in linearized form

$$\partial_s \widetilde{W} = \mathcal{L}_W(\widetilde{W}, \widetilde{Z}) + \mathcal{F}_W \quad \text{and} \quad \partial_s \widetilde{Z} = \mathcal{L}_Z(\widetilde{W}, \widetilde{Z}) + \mathcal{F}_Z$$

where $(\mathcal{F}_W, \mathcal{F}_Z)$ include nonlinear and dissipative terms.

- ▶ We study the linearized operator $\mathcal{L} = (\mathcal{L}_W, \mathcal{L}_Z)$
- ▶ It is also sometimes useful to write the operator in terms of self-similar velocity \mathcal{U} and sound speed \mathcal{S} , so that $\mathcal{L} = (\mathcal{L}_U, \mathcal{L}_S)$

Strategy for linear stability

1. Causality is used to cut-off the linearized equation on a neighborhood of the backwards acoustic cone ($|\zeta| \leq 1$). Redefine \mathcal{L} using cut-offs.
2. Show that for large m and small $\delta_g > 0$, \mathcal{L} decomposes as

$$\mathcal{L} = A = A_0 - \delta_g + K$$

for some A_0 maximally dissipative on H_{Rad}^m and K is some compact.

3. Use (U, S) variables for high derivative energy estimates (dissipativity) and (W, Z) for low derivative arguments (maximality).
4. Lumer-Phillips theorem ensures exponential decay modulo a finite dimensional unstable space.

Theorem

Let $\delta > 0$, and T the (strongly continuous) semigroup generated by $A = A_0 - \delta + K$ where $A_0 : H \rightarrow H$ is a maximally dissipative operator and $K : H \rightarrow H$ is compact. Then, there are finitely many eigenvalues λ_i with $\operatorname{Re}(\lambda_i) \geq 0$.

Let $\psi_i \in H$ be the corresponding eigenfunctions, let V be the finite dimensional space $V = \operatorname{span}(\psi_i)$. There exists $U \subset H$ such that U, V are invariant spaces of A and $H = U \oplus V$. Moreover

$$\|T(t)X\| \leq Ce^{-\delta t/2} \|X\|$$

for all $X \in U$

Dissipativity

Aim is to show that for (U, S) in a finite co-dim subspace of $H_{\text{Rad}}^m[0, 2]$

$$\langle \mathcal{L}(U, S), (U, S) \rangle_{H^m} \leq - \|(U, S)\|_{H^m}^2$$

- ▶ Highest order terms cancel.
- ▶ Left over with m th order terms have a good sign, if m is large.
- ▶ Projecting away low frequencies controls lower order terms.

Nonlinear stability argument

1. Control on unstable modes of truncated equations \implies global control on low order derivatives.
2. Global control on low order derivatives \implies global control on high order derivatives.
3. Global control on high order derivatives \implies dissipation can be treated as an decaying forcing term for the linearized problem.
4. Topological argument closes the argument, leading to a global solution of the self-similar equation.

Topological argument

Let $\{\psi_i\}_{i=1,\dots,N}$ be a basis for the **unstable manifold** and

$$\kappa_i = \langle (\tilde{U}, \tilde{S}), \psi_i \rangle$$

be the unstable modes of a solution. Let $\mathcal{R}(s) = B_{\mathbb{R}^N}(0, e^{-\delta s})$. If $\kappa \in \partial\mathcal{R}$, we can show that κ leaves \mathcal{R} immediately.

Consider solutions with initial unstable modes $\kappa(s_0) \in \mathcal{R}(s_0)$. Suppose all such initial data leave \mathcal{R} in finite time. This would imply (after rescaling) the existence of a continuous map from $B_{\mathbb{R}^N}(0, 1)$ to \mathcal{S}^{N-1} , which leads to a contradiction.

Computer assisted arguments

Computer assisted interval arithmetic is used to prove the positivity of certain quantities: e.g. positivity of a polynomial over finite interval (barrier arguments) or the sign of a Taylor coefficient.

We define an arithmetic such that for intervals X, Y and $x \in X, y \in Y$

$$x \star y \in X \star Y.$$

for a given operator \star . E.g.

$$[\underline{x}, \bar{x}] + [\underline{y}, \bar{y}] = [\underline{x} + \underline{y}, \bar{x} + \bar{y}]$$

$$[\underline{x}, \bar{x}] \times [\underline{y}, \bar{y}] = [\min\{\underline{x}\underline{y}, \underline{x}\bar{y}, \bar{x}\underline{y}, \bar{x}\bar{y}\}, \max\{\underline{x}\underline{y}, \underline{x}\bar{y}, \bar{x}\underline{y}, \bar{x}\bar{y}\}].$$

Thank you!